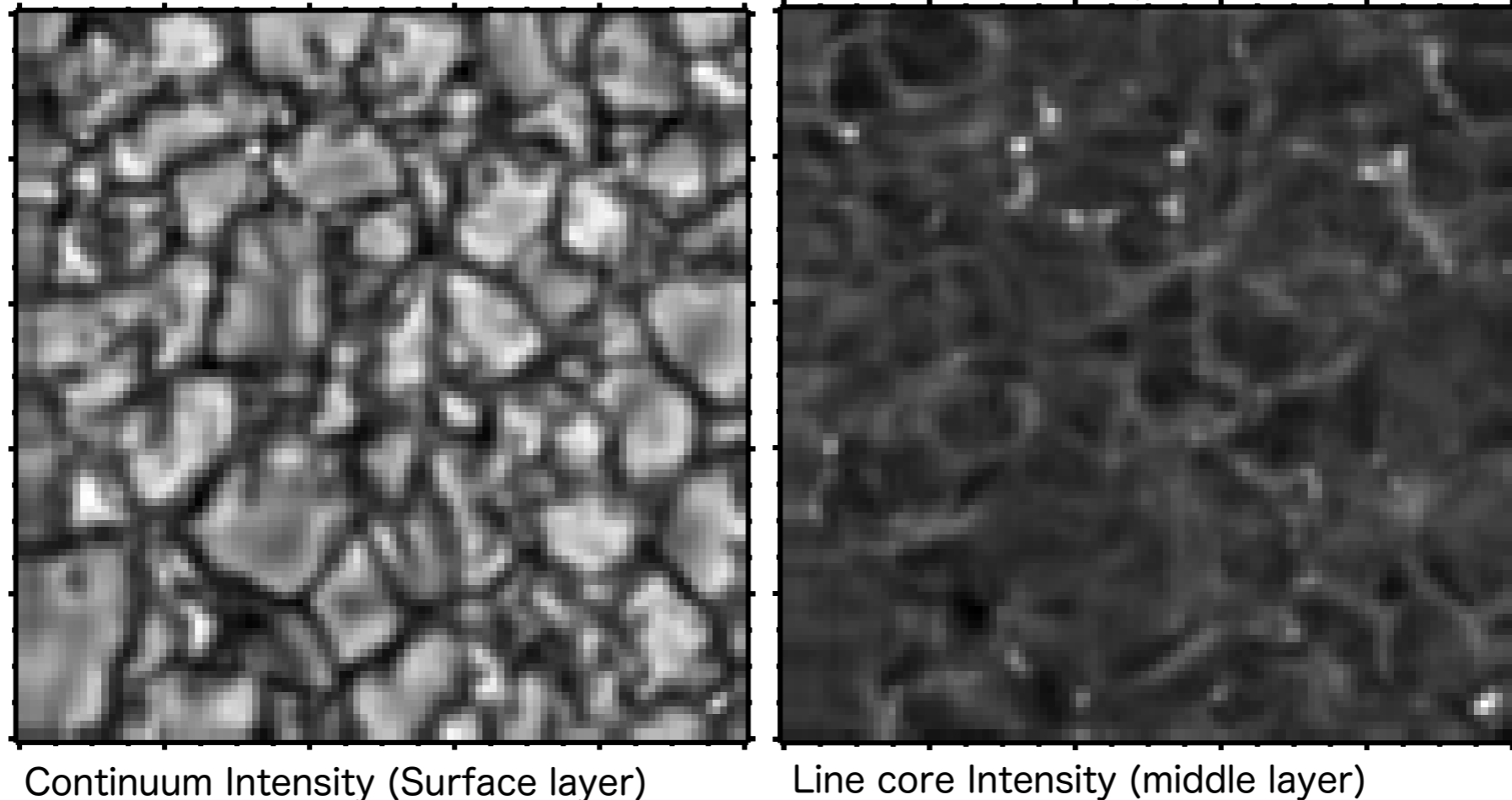


Study of the relation between the reversed granulation and the gas compression/expansion

taken by Hinode-SOT/SP



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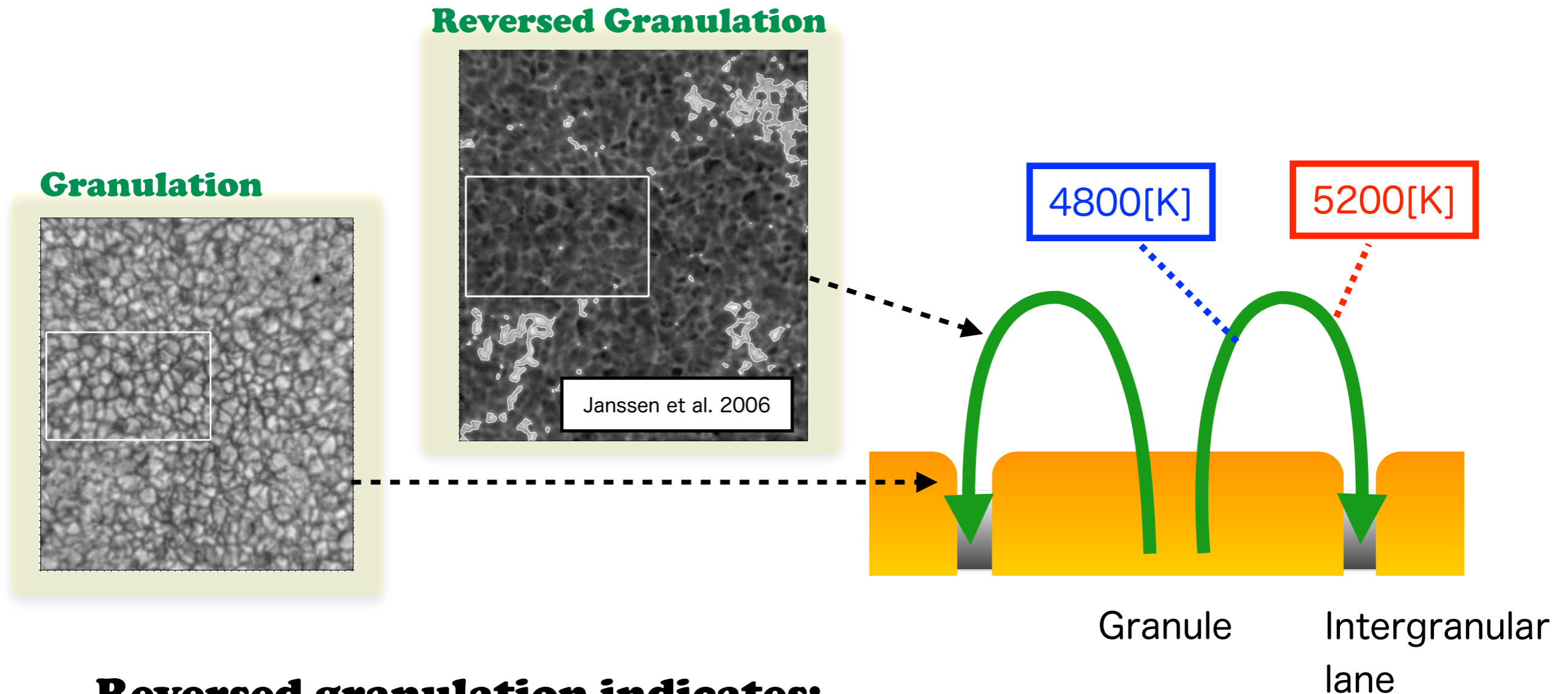
4. Institute of Space and Astronautical Science / Japan Aerospace Exploration Agency

What is Reversed Granulation?

Intensity pattern at the middle layer of the photosphere (height > ~130km)
is opposite to that of the surface granulation

Ruiz Cobo et al. 1996

Faurobert et al. 2013



Reversed granulation indicates:

Gas material, when moving from granule to intergranular lane, is somehow heated up.

Several candidates to explain the reversed granulation

Magnetic field (Unlikely?)

Leeanarrrts et al. 2005

- It could be seen in internetwork region
- Non-magnetized numerical simulation also reproduces it

p-mode, internal gravity wave

de Wijn 2009 Rutten 2004

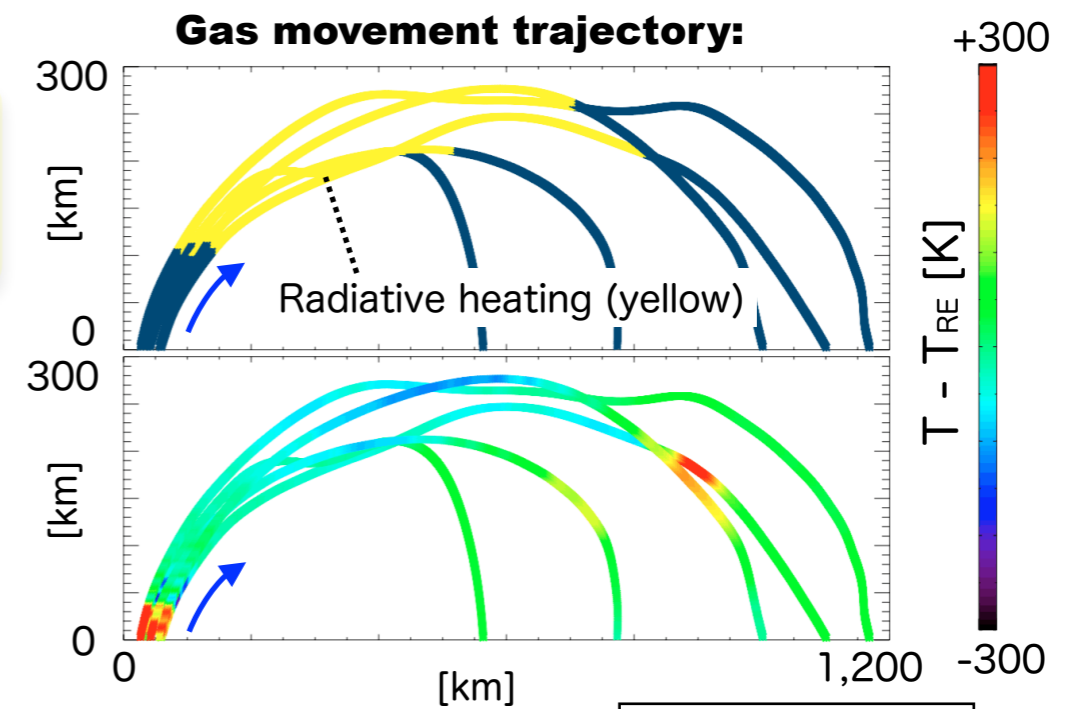
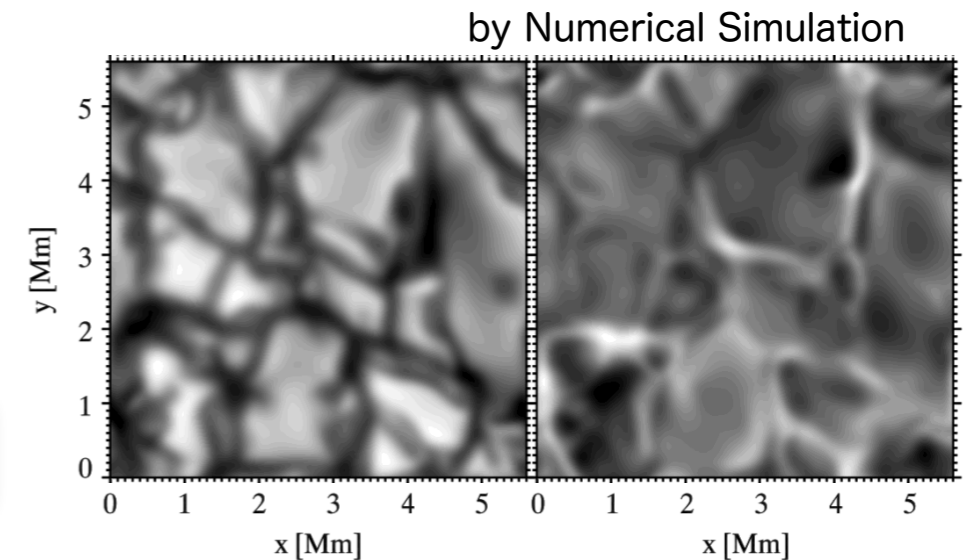
- Observation shows they are cospatial with the reversed granulation pattern

Gas-compression, Radiative heating

Nordlund et al. 1985

- Gas is compressed in intergranular lane by surrounding granular flow
- Gas in intergranular lane at the middle height is heated by radiation

$$\frac{D \ln T}{Dt} = \underbrace{-(\gamma_3 - 1) \nabla \cdot \mathbf{v}}_{\text{compression/expansion}} + \underbrace{\left(\frac{\partial \ln T}{\partial s} \right)_e \frac{Q_{\text{rad}}}{\rho T}}_{\text{Radiation}}$$



Cheung et al. 2007

➔ Several ideas are proposed but not conclusive, due to mainly lacking their observational proofs

Observational Studies in the Past

Balthasar et al. 1990

Correlation coefficient between I_{cont} and I_{core} (typically $-0.2 \sim -0.3$)

Poor resolution

Rutten et al. 2004

Janssen et al. 2006

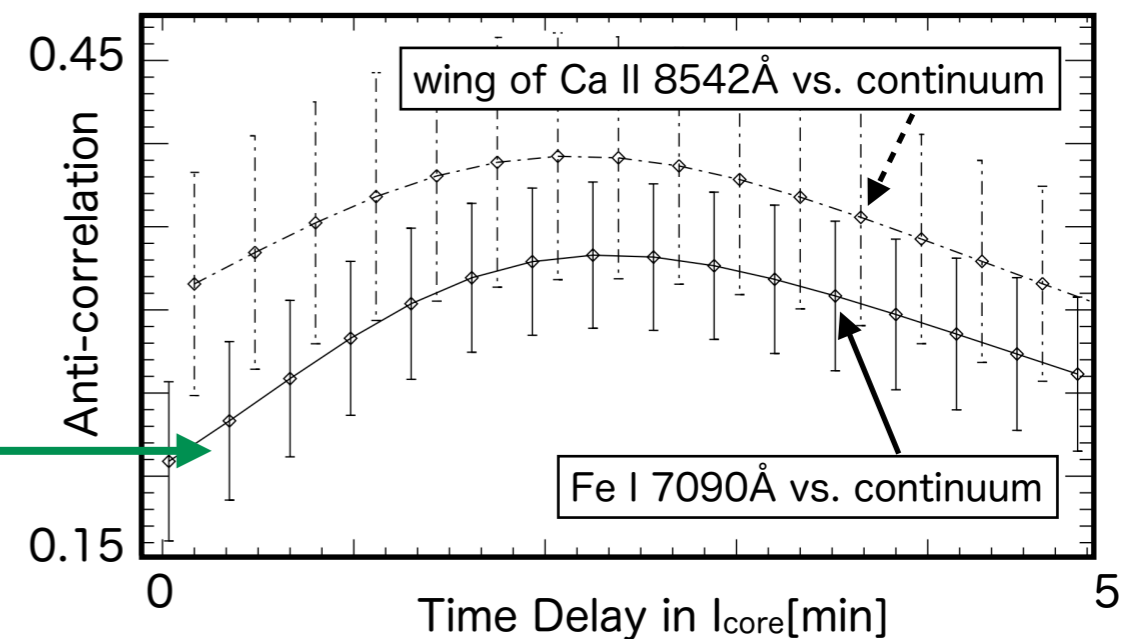
→ Better Correlation: $-0.4 @ 1.5''$

Time delay of 2 min in I_{core} behind I_{cont}

→ Better Correlation: -0.5

※ upflow with 2 km/s takes 2 min to reach 120 km

➔ Clearly somehow related to gas convection



Issue

No reports for the relation between horizontal flow field (gas compression/expansion) and the reversed granulation, so far

This study ➔ Doppler analysis at the solar Limb

Observation and Analysis

Hinode/SP

Spatial resolution: 0.3"

Spatial sampling: 0.15"x0.16"

Region: [0", 760"]*

*In the LOS, 63% of horizontal flow is reflected

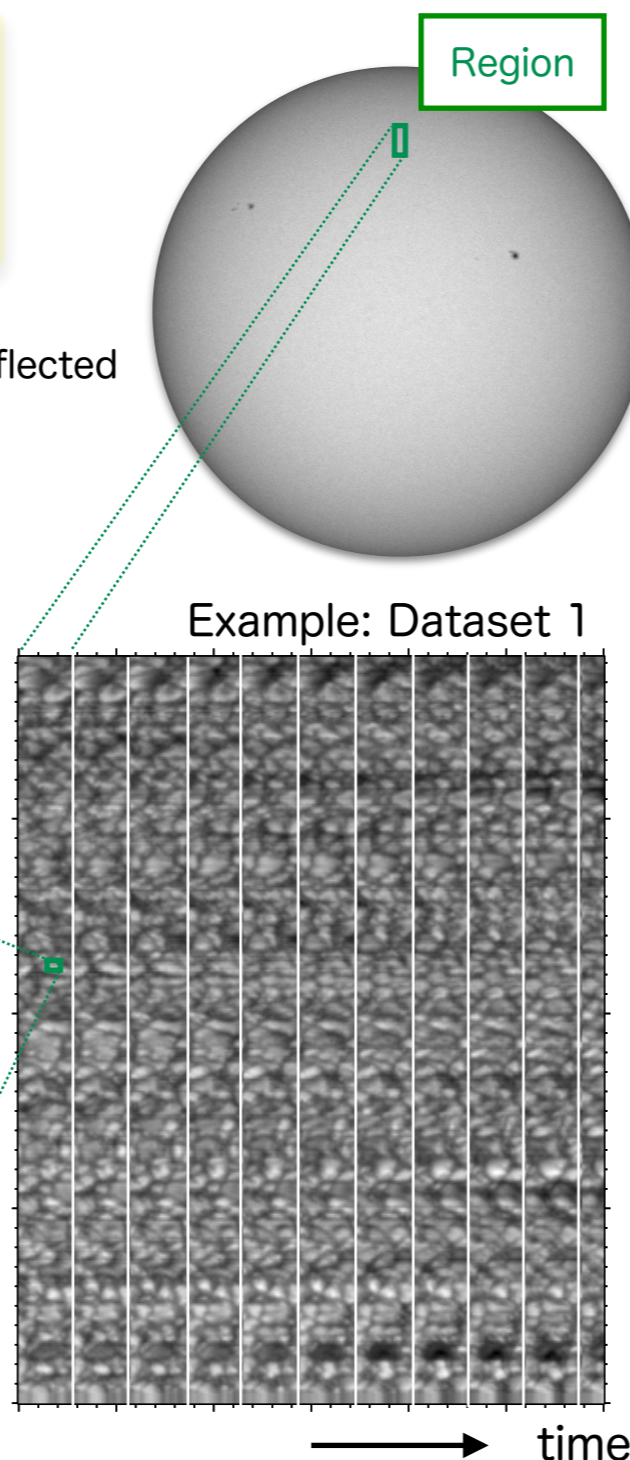
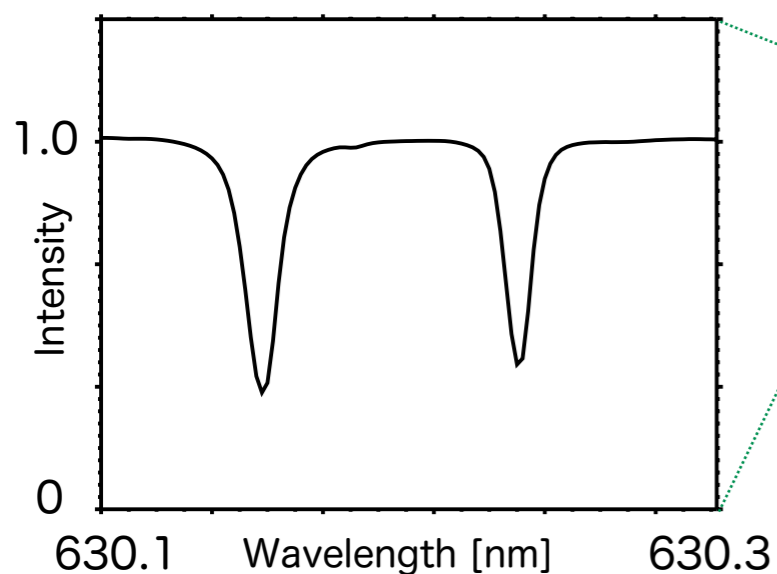
Dataset 1

FOV: 4.5" x 61"

Cadence: 30 sec

Dataset 2

FOV: 90" x 162"

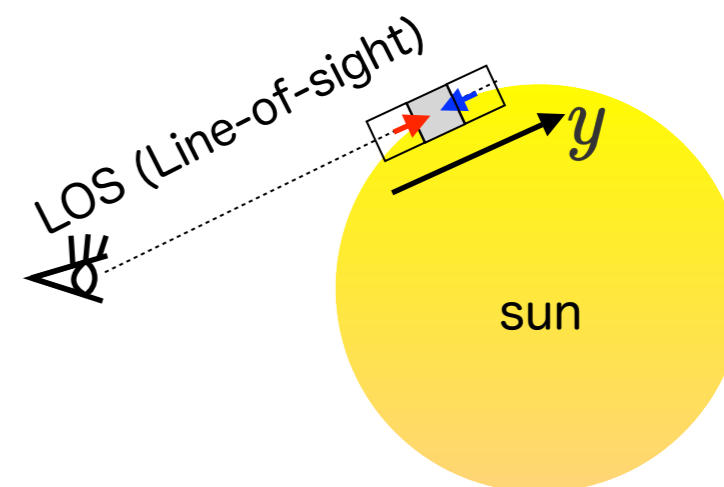
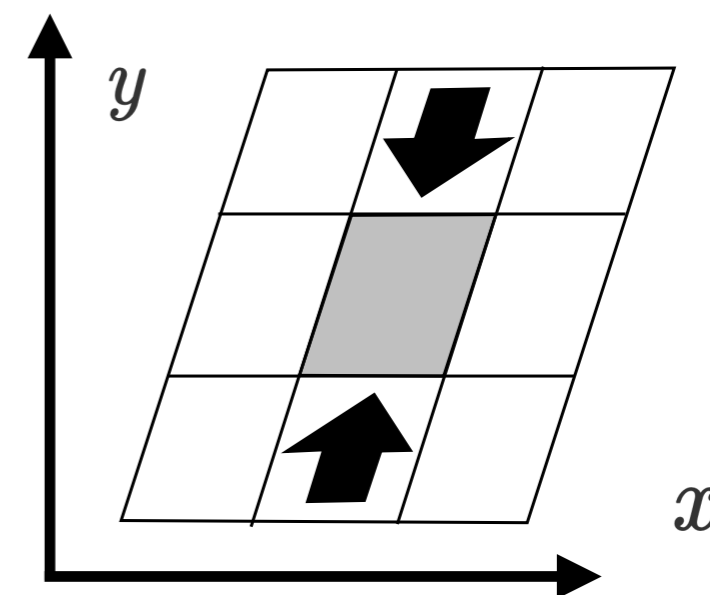


*Deconvolution analysis is applied

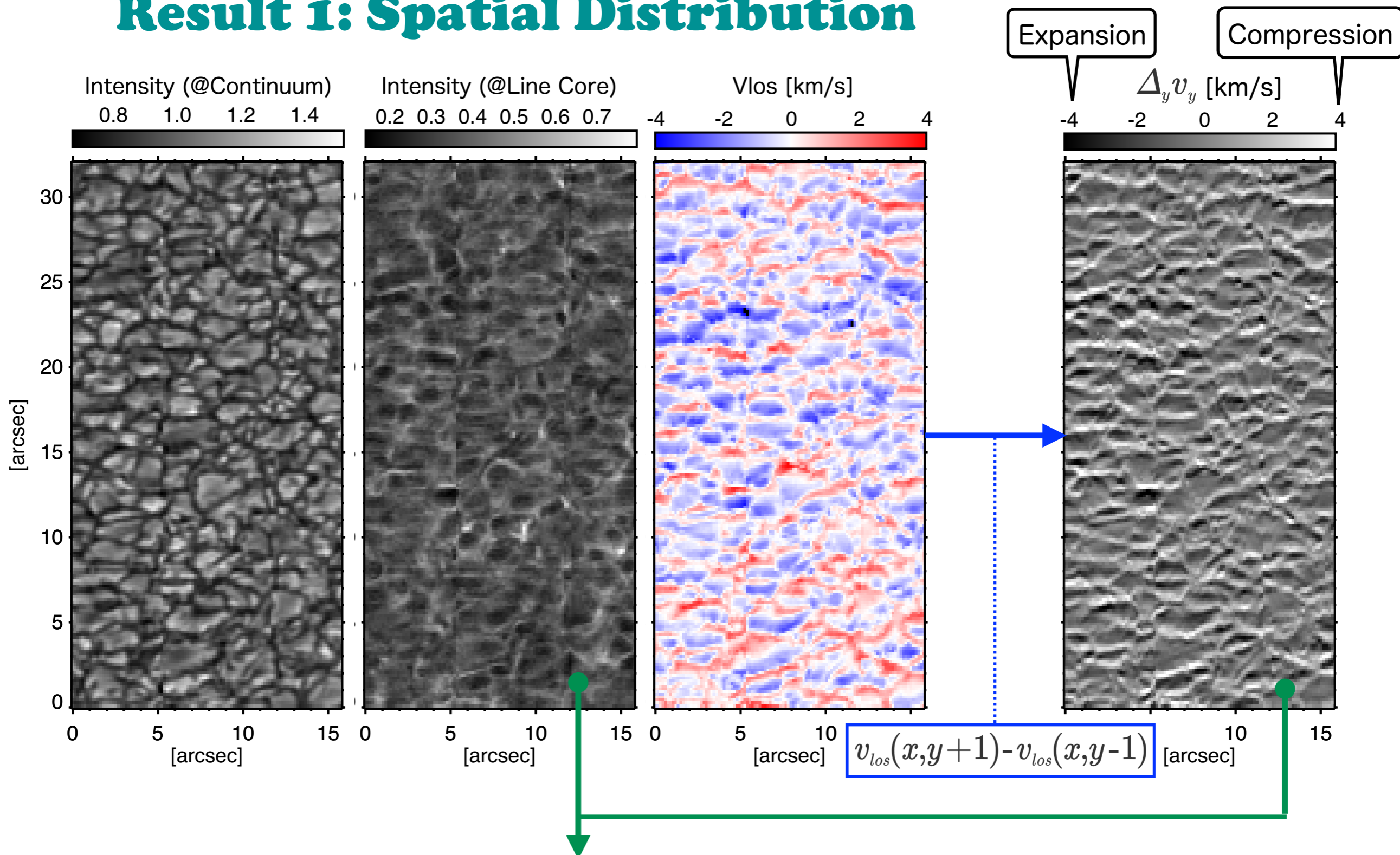
Oba et al. 2017

Evaluation for Expansion/Compression

$$\Delta_y v_y(x,y) \text{ is introduced} \\ = v_{los}(x,y+1) - v_{los}(x,y-1)$$



Result 1: Spatial Distribution

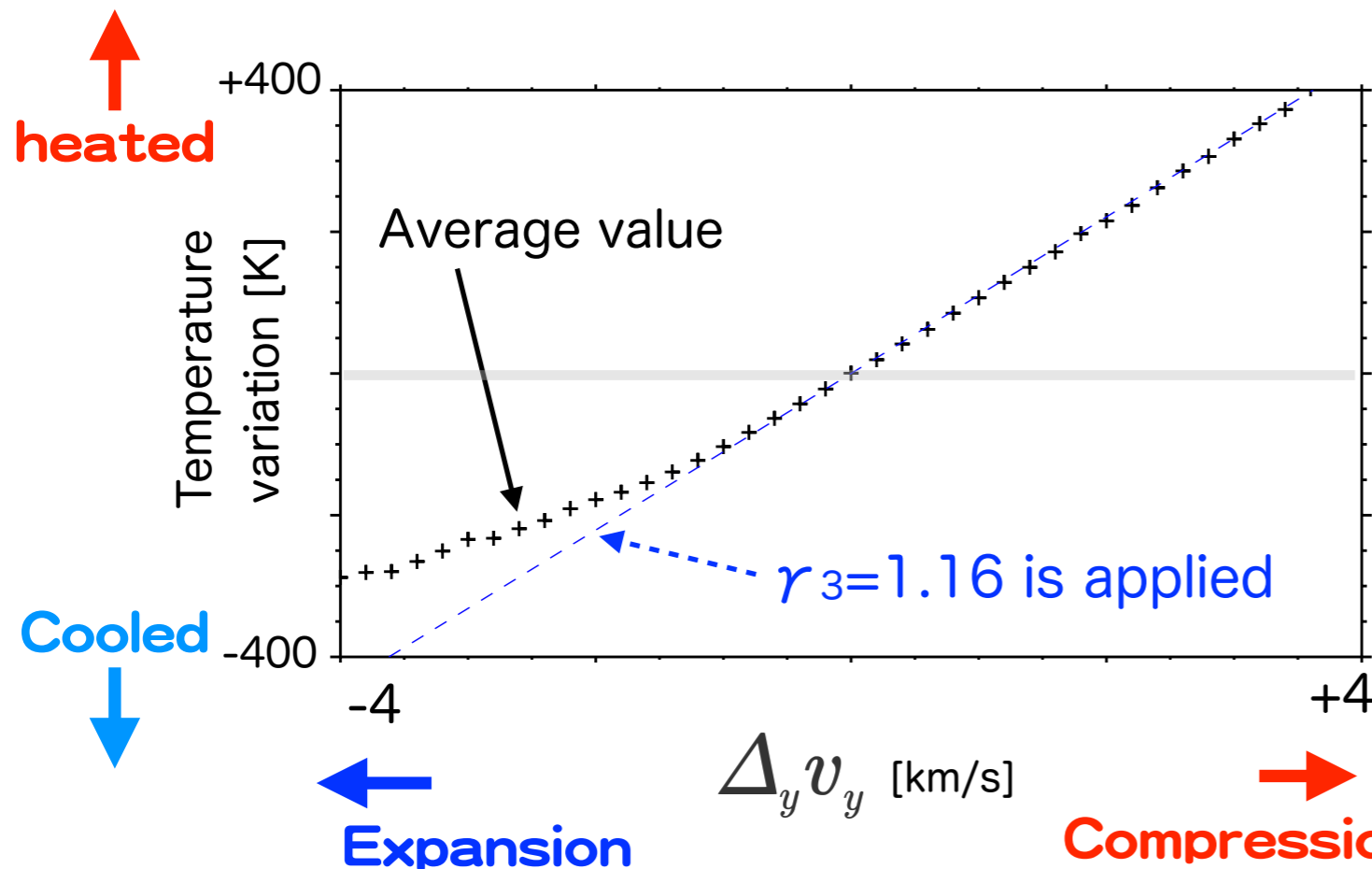


Correlation Coefficient

Intensity (@Line core) vs. Gas compression/expansion: 0.64

Result 2: Average of Temperature variation as a function gas compression/expansion

※ Temperature: determined from intensity via the Planck function



Cheung et al. 2007

Mihalas et al. 1985

※ ignore the radiation term

$$\frac{D \ln T}{Dt} = -(\gamma_3 - 1) \nabla \cdot \mathbf{v} \propto \Delta_y v_y$$

↑
Adiabatic exponent

Temperature variation:
proportional to gas compression/expansion

Temperature variation at the middle photosphere

It is determined by the act of compression/expansion

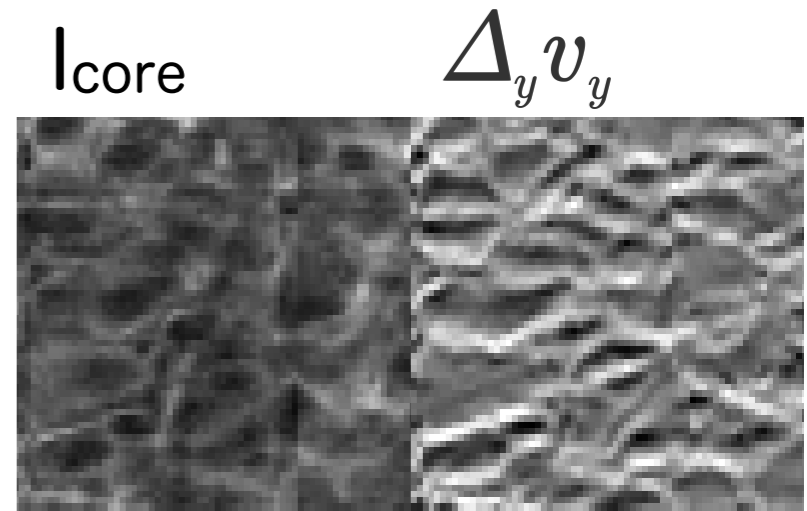
Compression increases the temperature, while expansion does not well decrease the temperature

Discussions

For result 1: Compression/Expansion vs. l_{core}

Strong correlation: 0.64

Temperature variation at the middle layer could be qualitatively explained by the gas expansion/compression



Expansion@granule ← Overshoot with decreasing pressure
Compression@Intergranular lane ← Colliding with anti-directed flow

For result 2: Temperature variation due to gas compression / Expansion

Compression: $\gamma_3=1.16$, Expansion: $1 < \gamma_3 < 1.16$

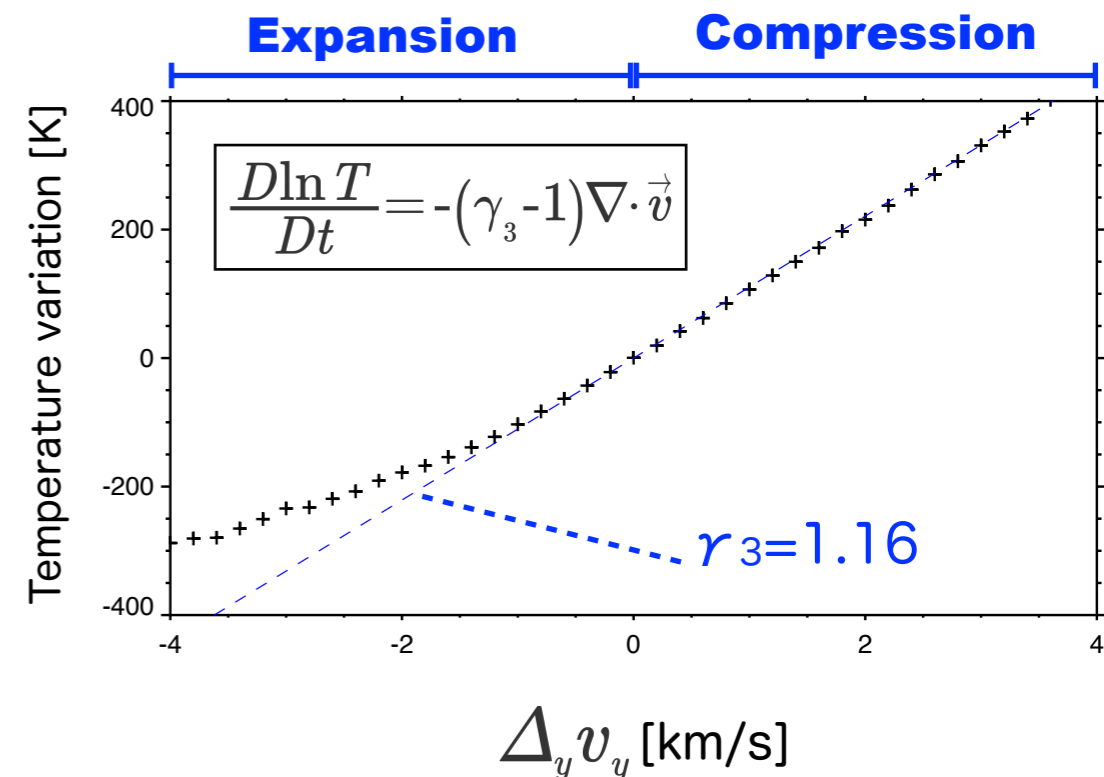
Compression increase the temperature,
 while expansion does not well decrease the temperature



- Radiative heating above granule contributes to heat partially

and/or Nordlund et al. 1985 Cheung et al. 2007

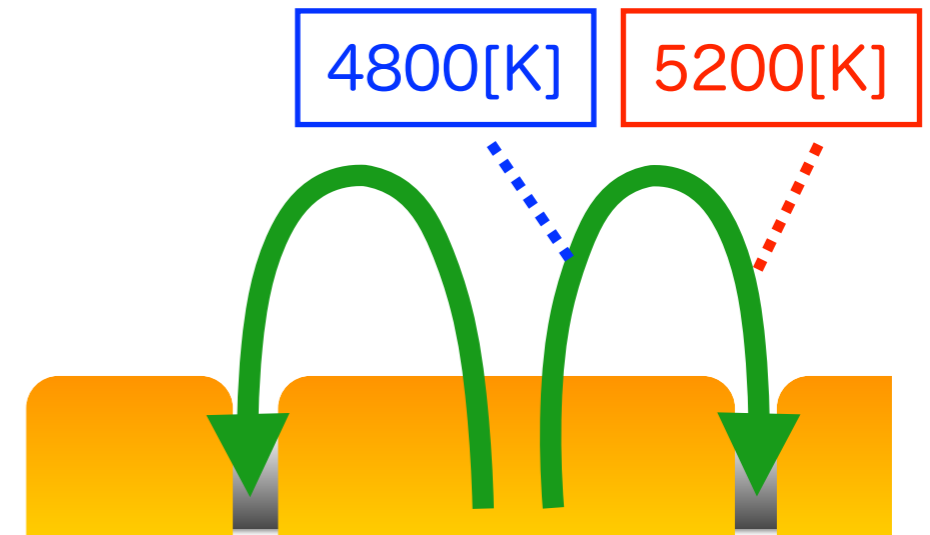
- Change of γ_3 in granules and intergranular lanes



Summary

Reserved granulation

Temperature enhancement when moving from a granule to an intergranular lane



Candidates Gas compression/Expansion, p-mode, internal gravity wave, ...

This study: Horizontal flow field derived from observations is used for the first time to evaluate the gas compression/expansion

➔ Observationally support the gas compression/expansion contributes to create the reserved granulation

※Radiation may partially heat the gas above granules

Future works

- Quantitative evaluation of γ_3 in granule and intergranular lane using numerical simulation
- Multiple Doppler analysis in coordination with the Solar Orbiter

